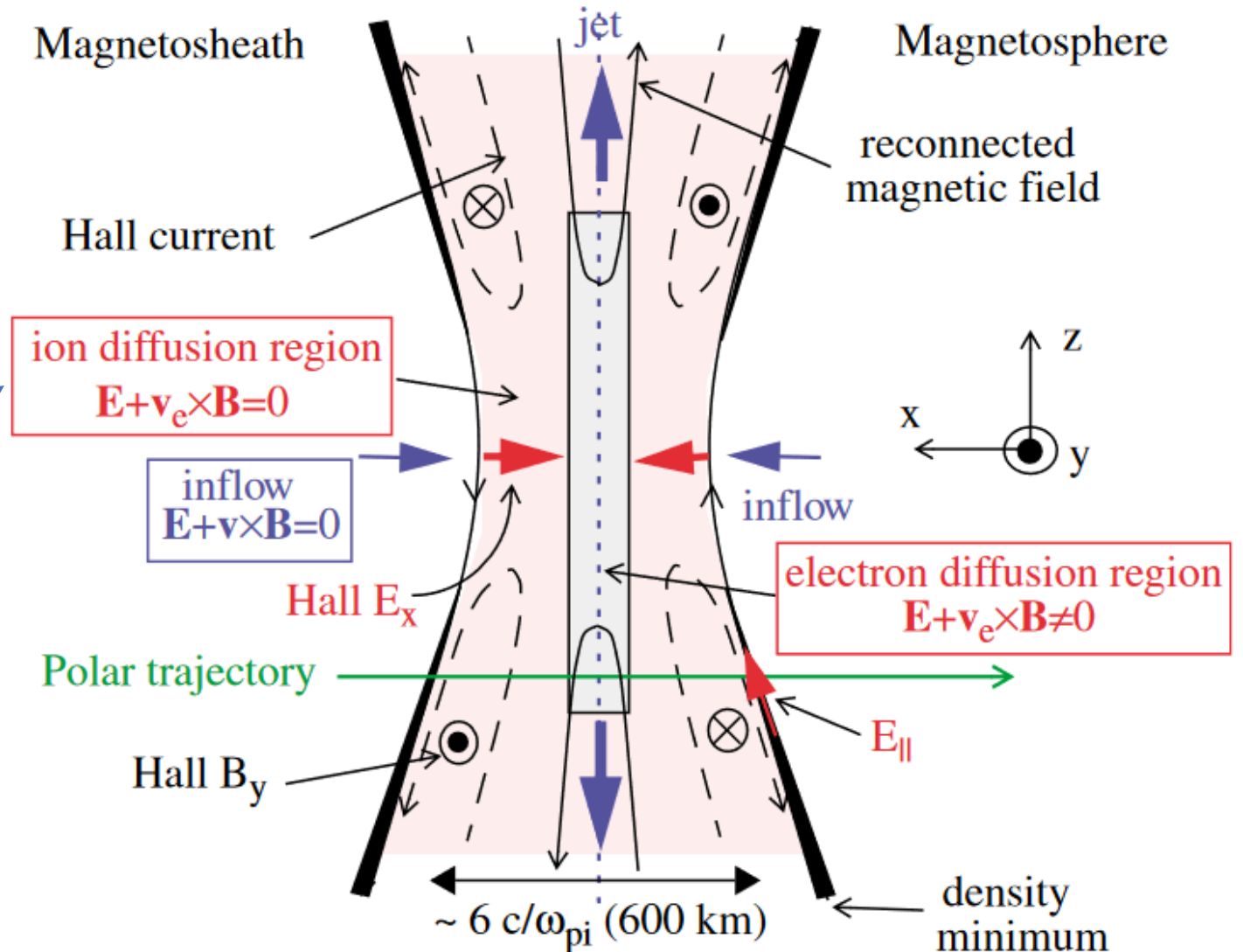


Hall Physics Above the Lunar Surface and Implications for Magnetic Reconnection

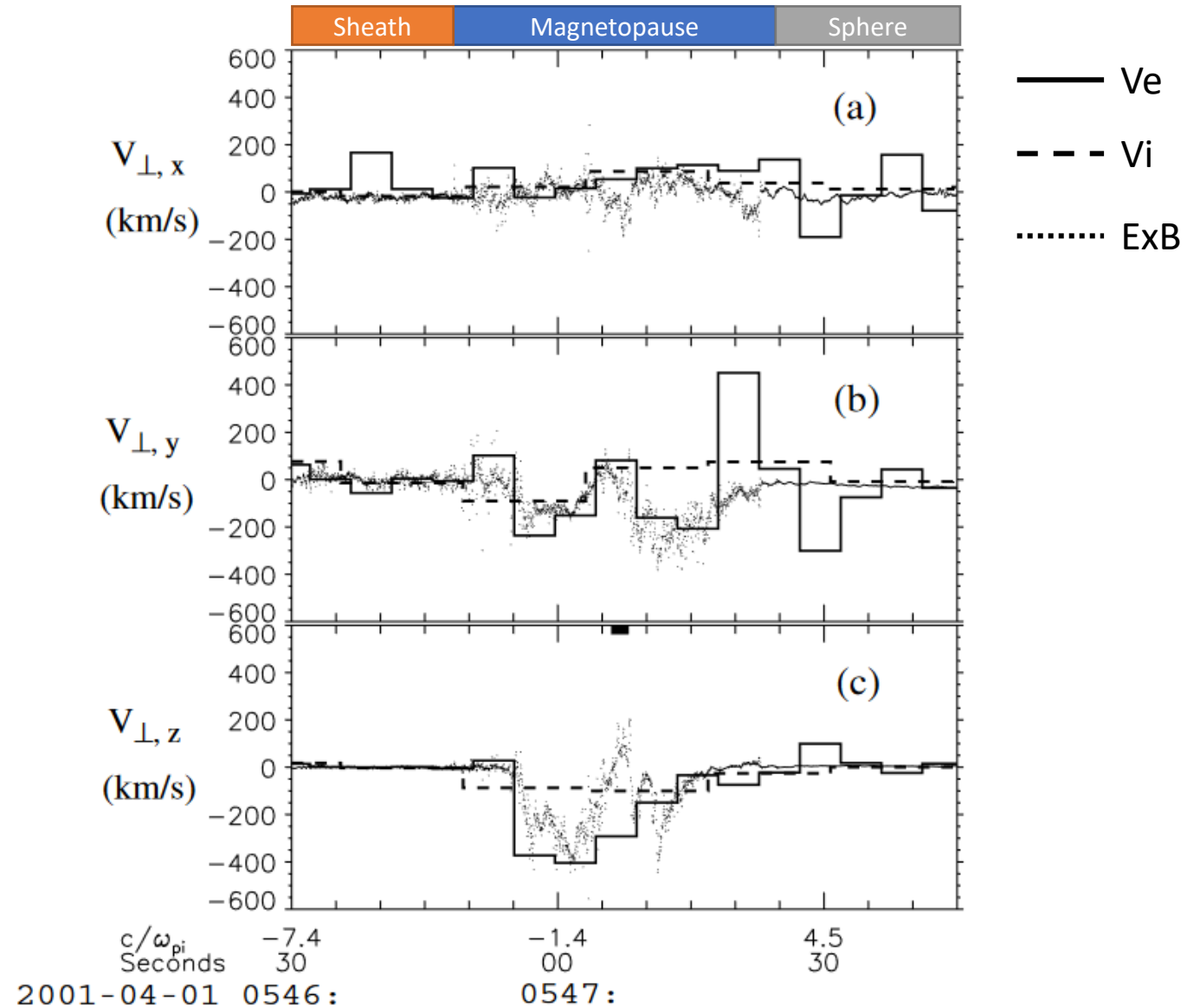
R. P. Sawyer, J. S. Halekas, J. W. Bonnell², L. J. Chen³, J. McFadden², K. H. Glassmeier⁴, Y. Harada⁵

Standard Ion-Coupled reconnection picture, showing Hall E-field structure and diffusion regions.



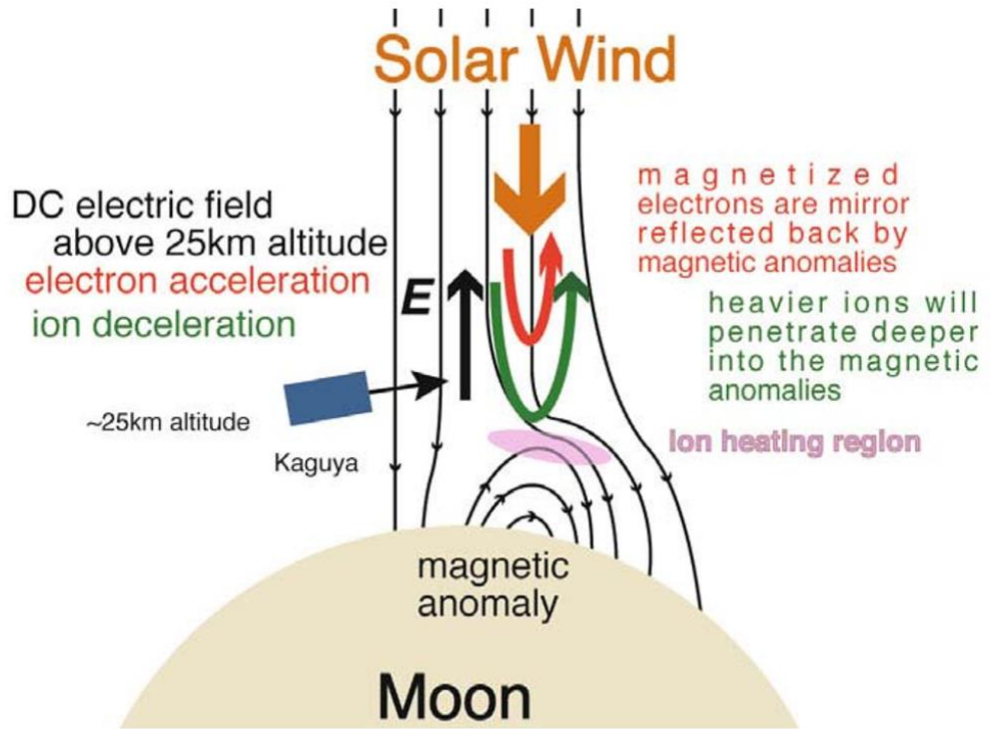
Ion diffusion region encounter near the magnetopause by Polar

Polar magnetopause crossing showing evidence of demagnetized ions and magnetized electrons.

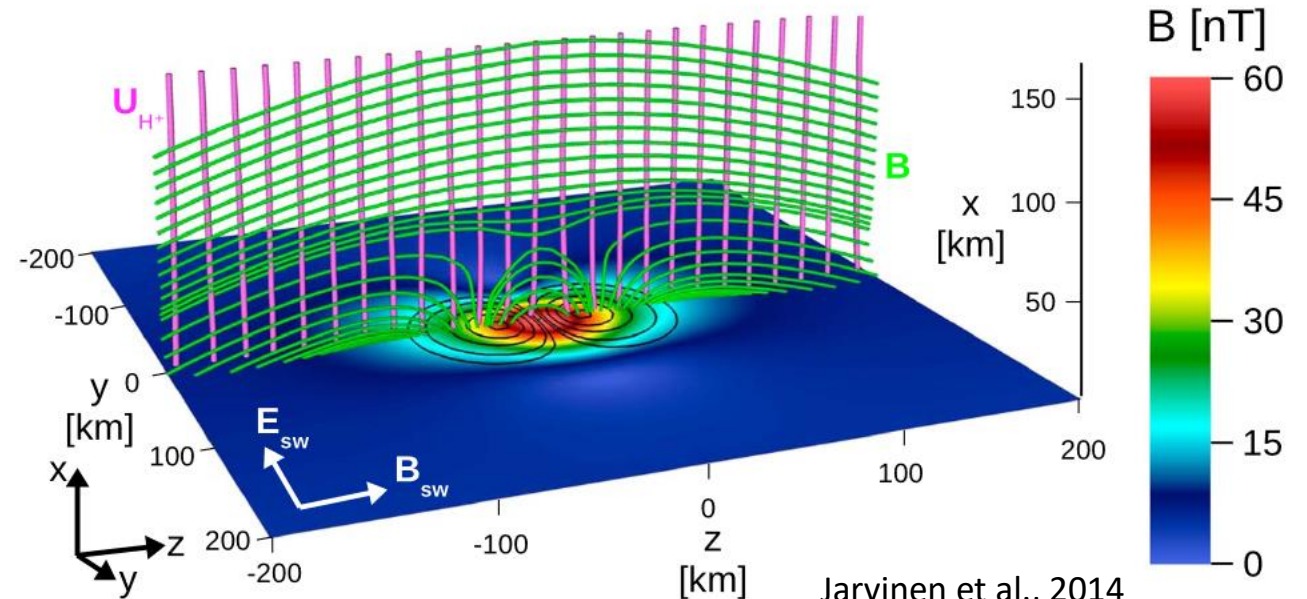
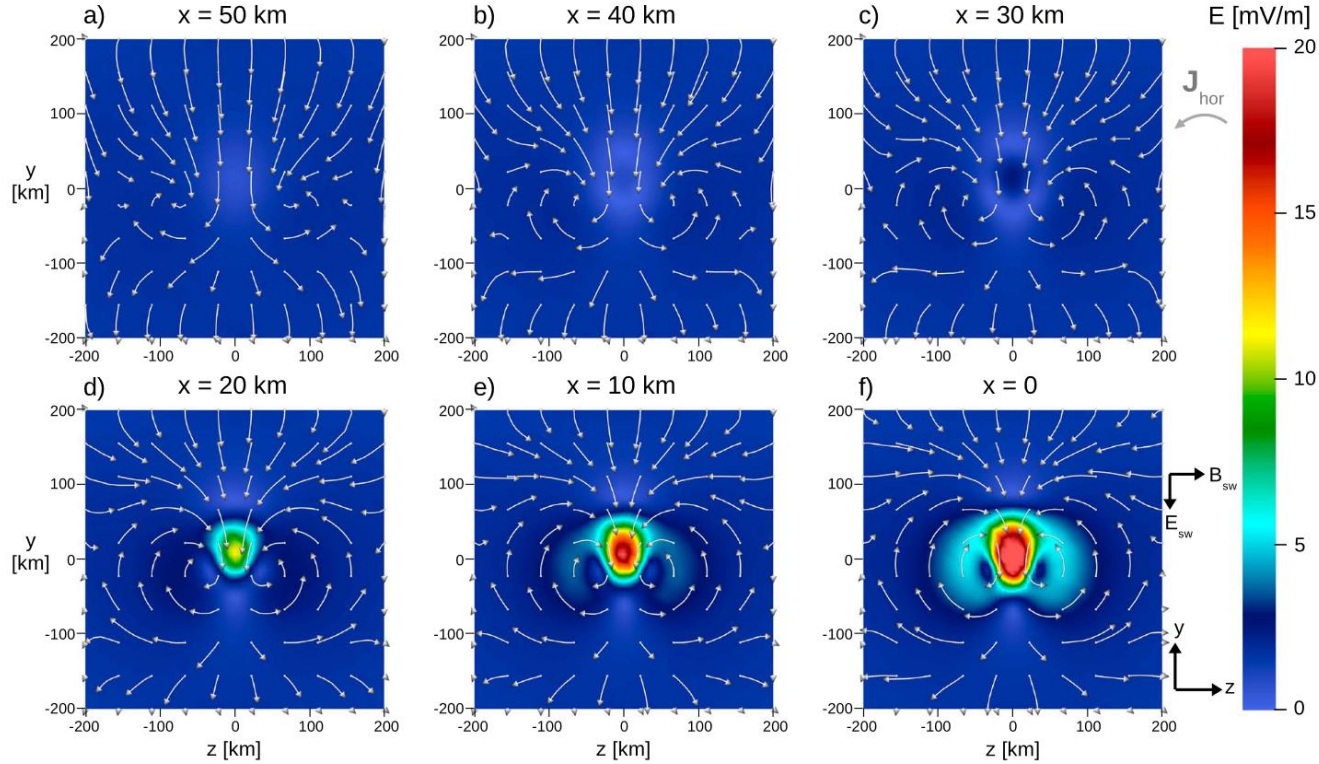


How does the near lunar surface environment compare?

Results from Saito et al., 2012 consistent with Hall MHD, and simulations from Jarvinen et al., 2014 show a similar field structure.

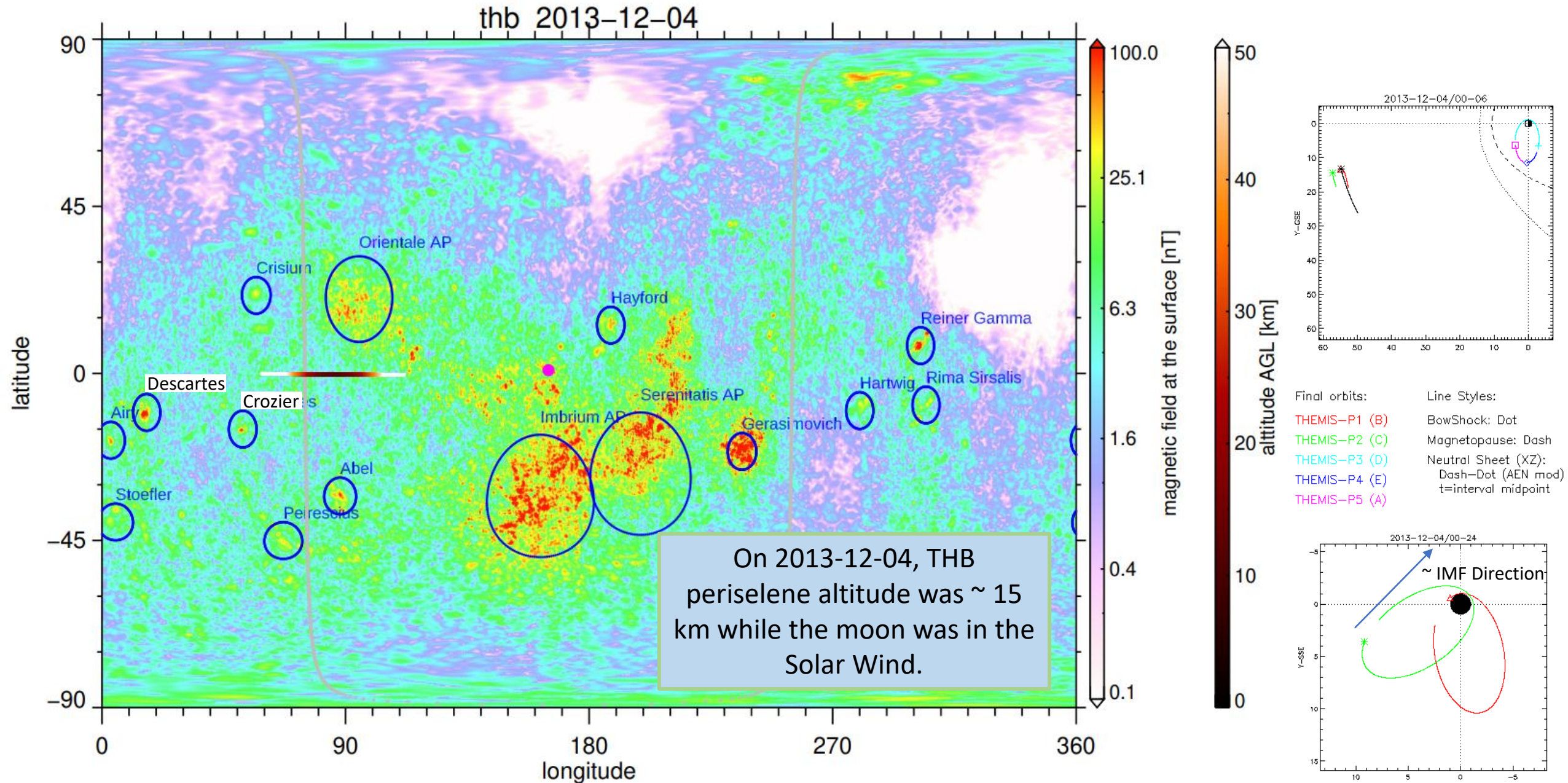


Saito et al., 2012

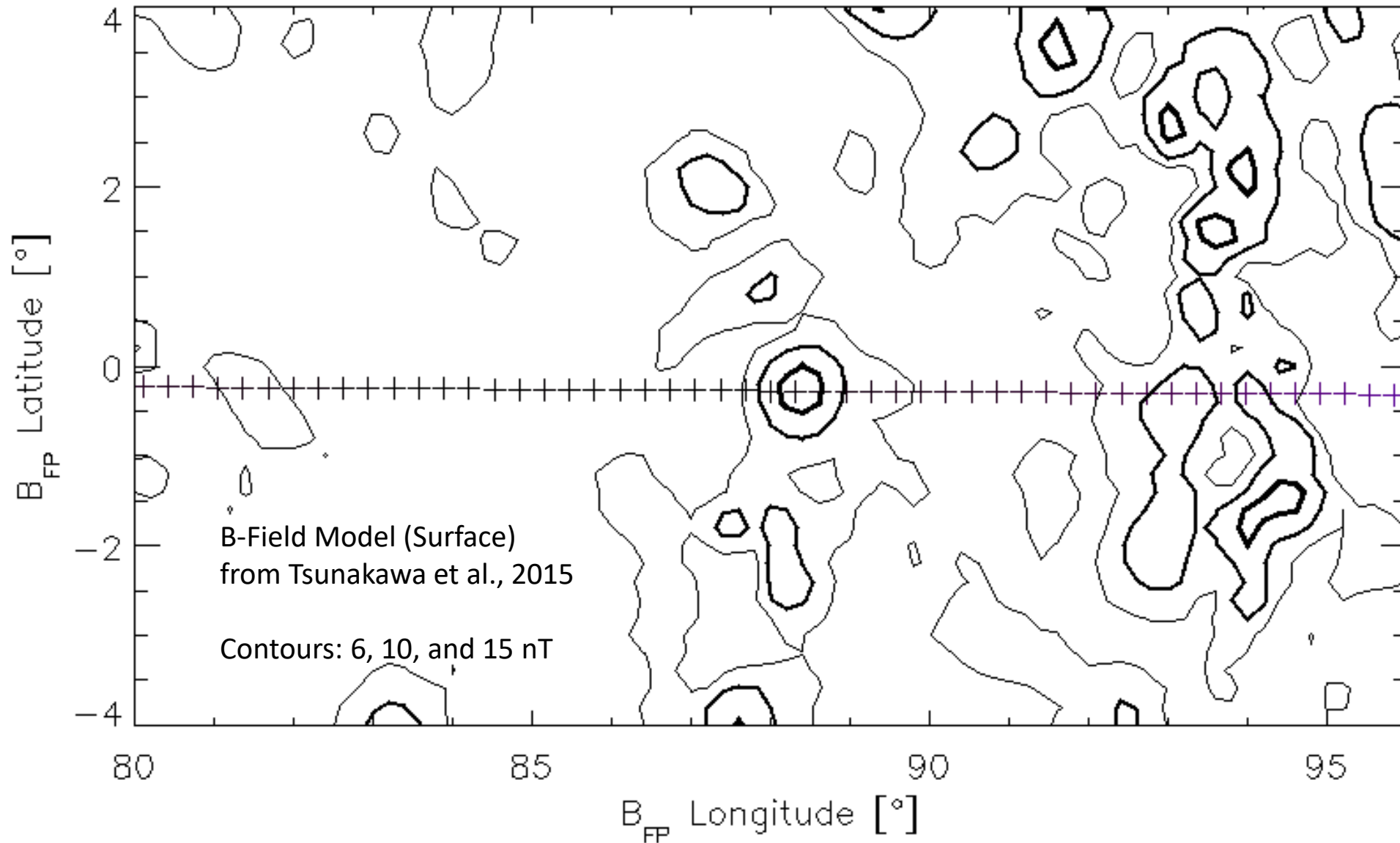


Jarvinen et al., 2014

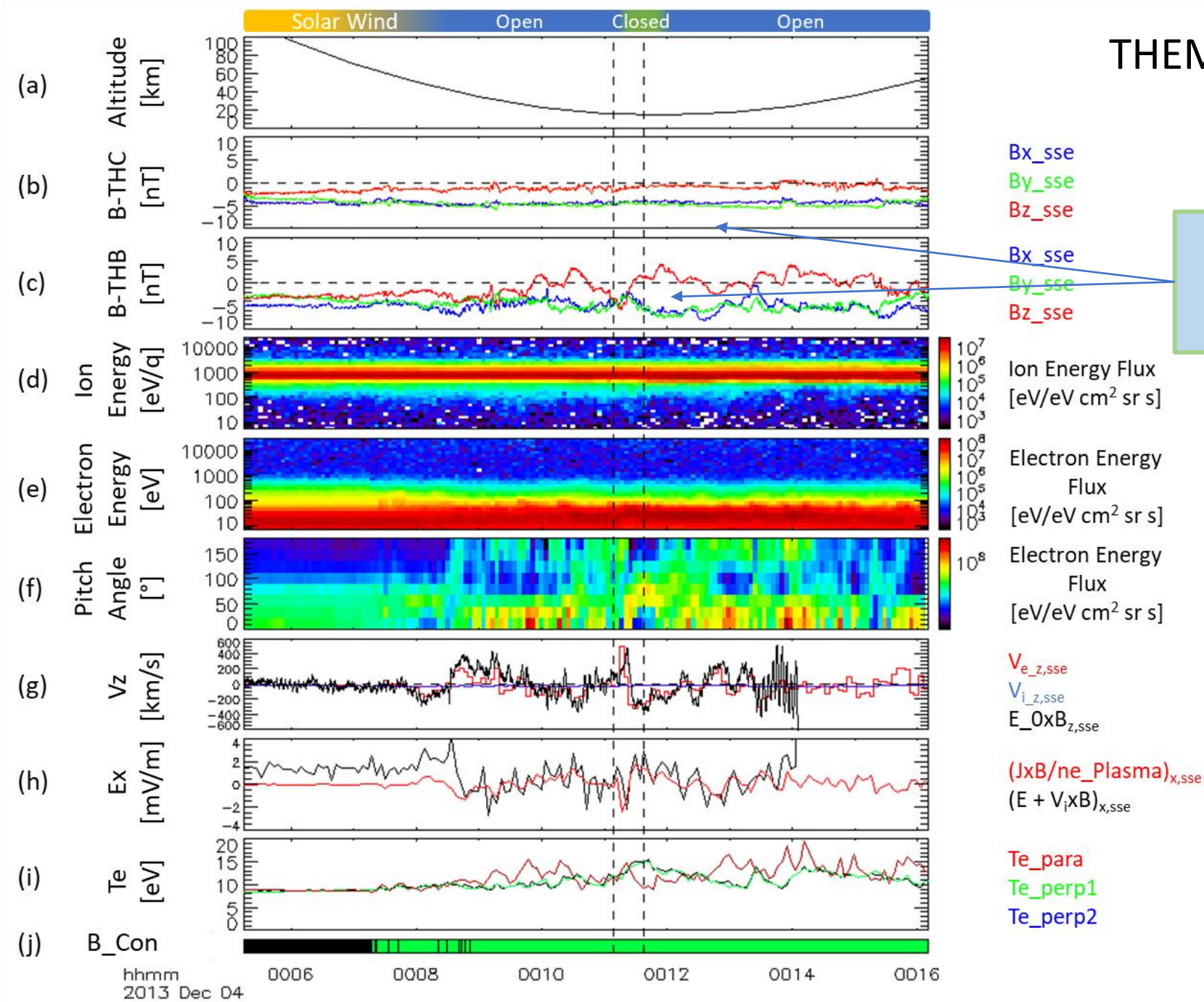
Where was THEMIS-ARTEMIS in its orbit?



Where was THEMIS-ARTEMIS in relation to the crustal fields on the surface?

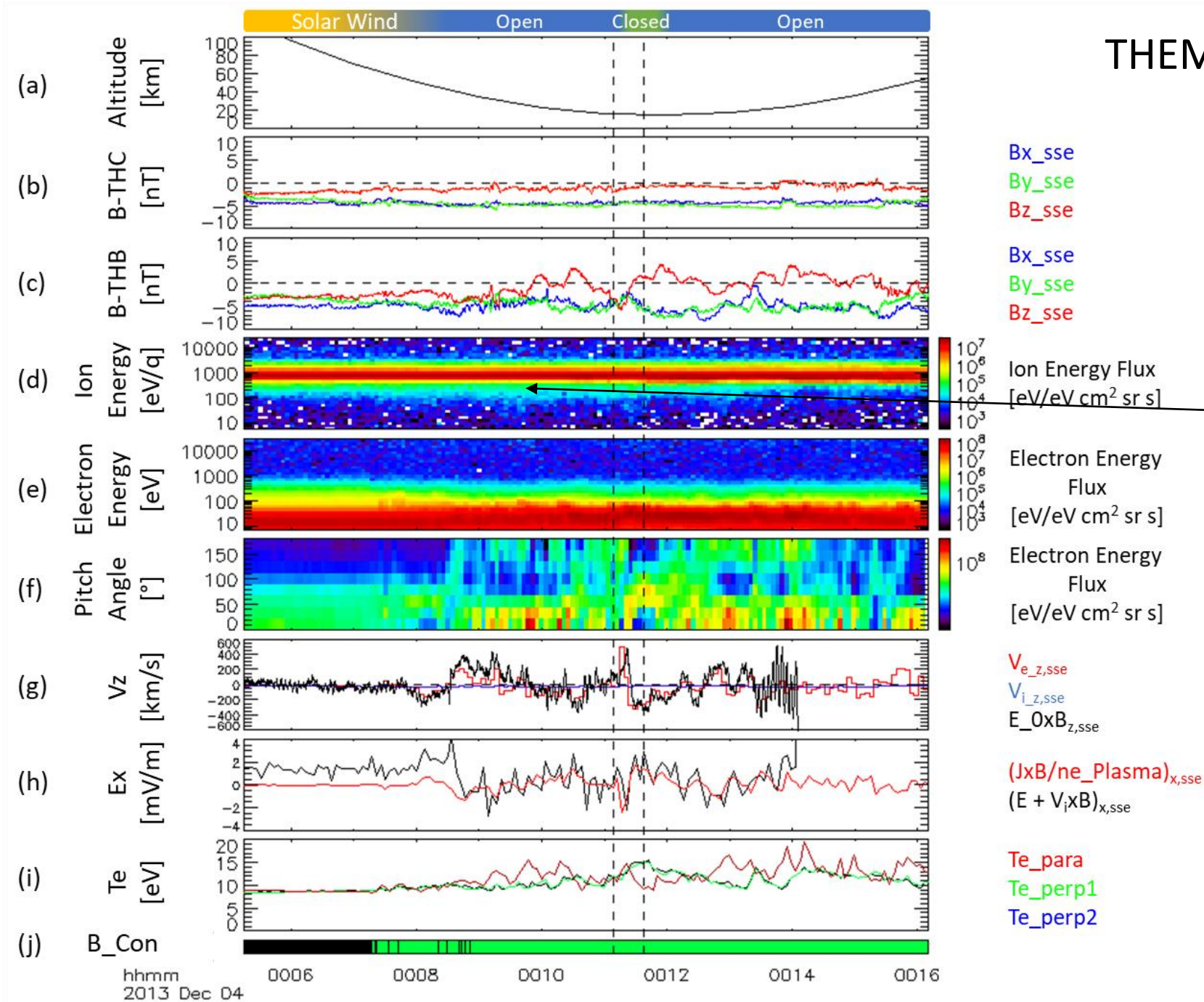


THEMIS-ARTEMIS observations of field line modification



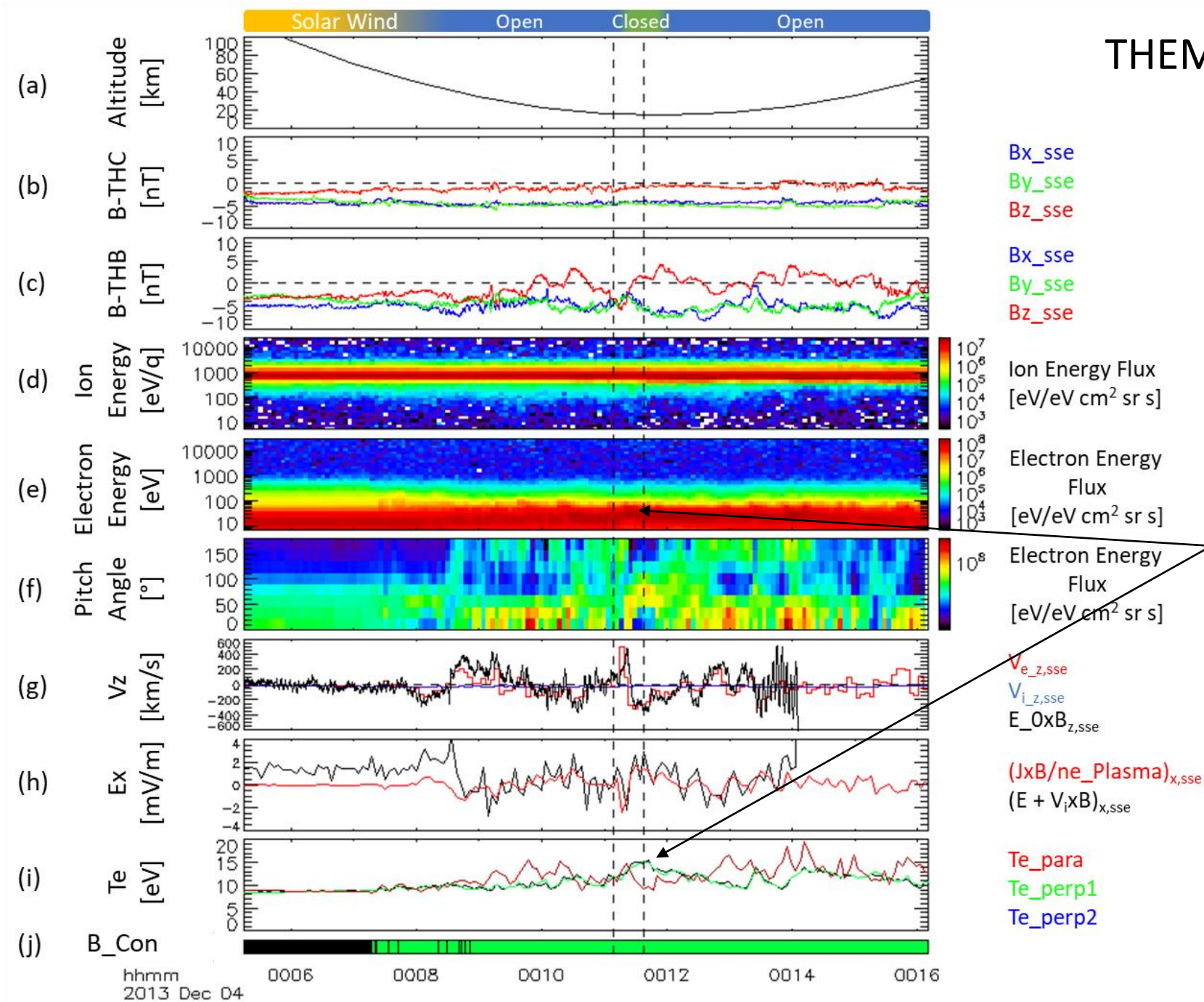
Clear indication of modification by crustal magnetic field in the magnetometer data.

THEMIS-ARTEMIS observations of solar wind ion reflection



Reflected ion signature

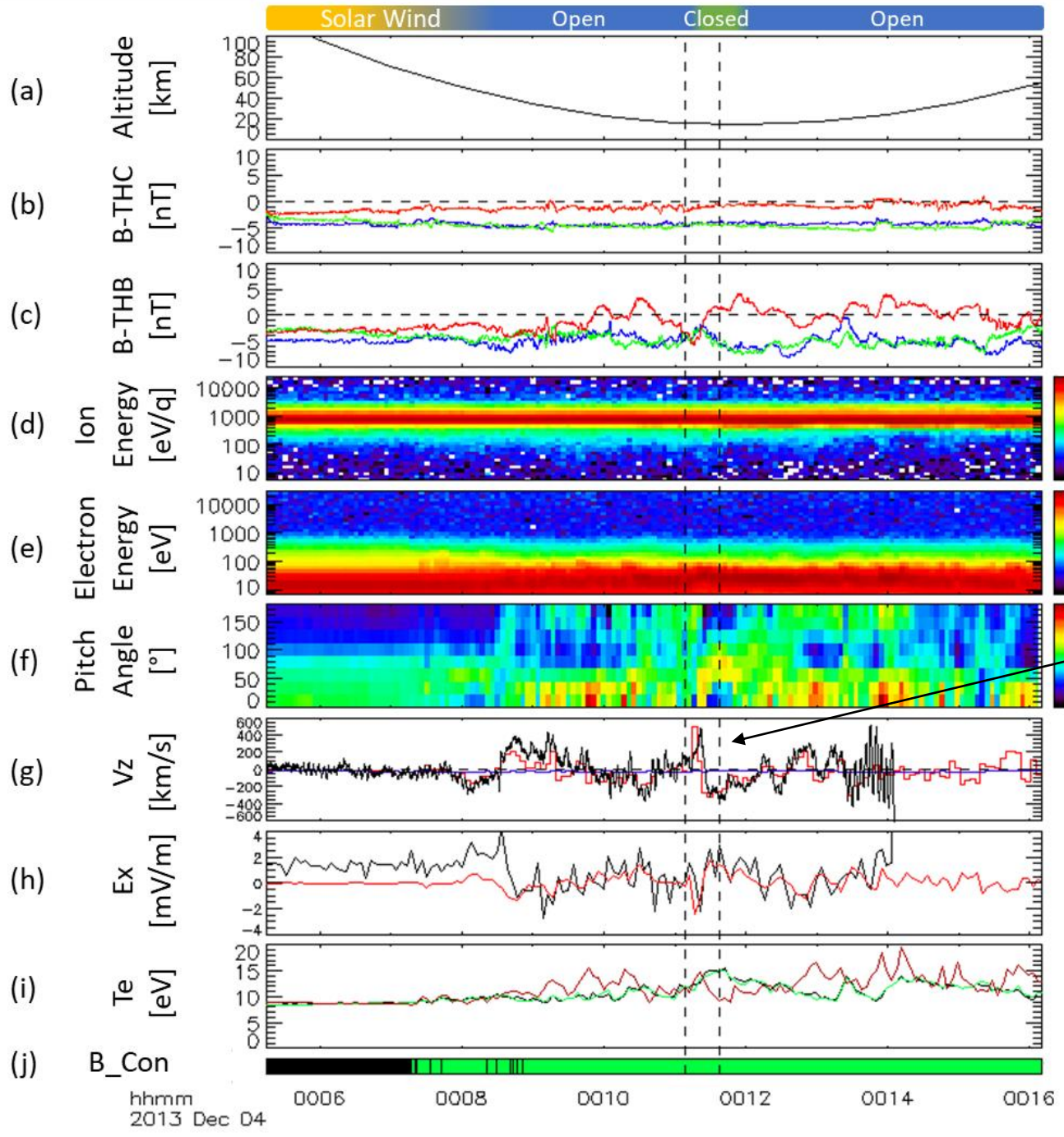
THEMIS-ARTEMIS observations of solar wind electron heating



Electron heating

Result 1:

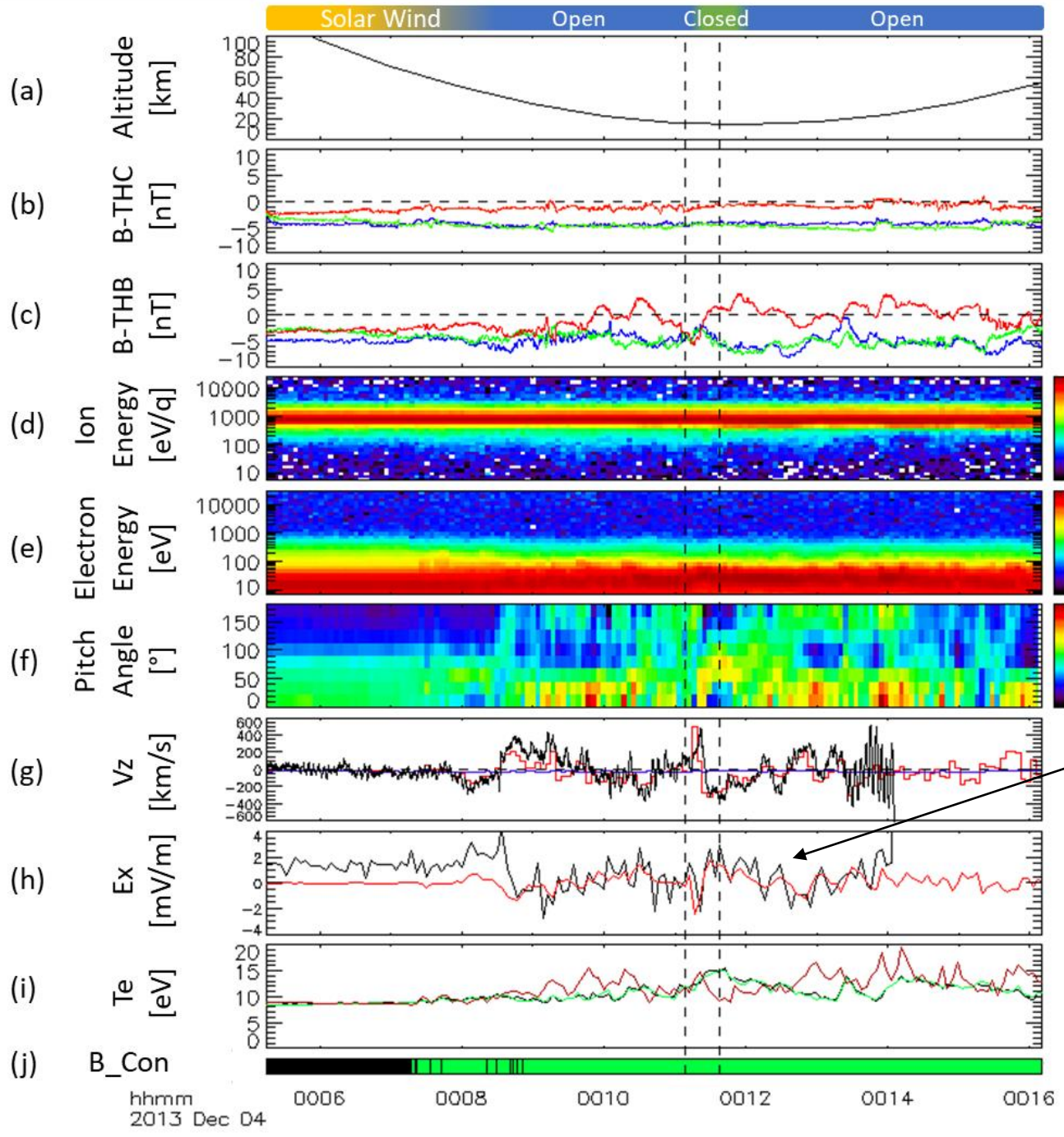
THEMIS-ARTEMIS observations of demagnetized ions and magnetized electrons



Agreement between V_{ez} and $E \times B_z$ indicating magnetized electrons with demagnetized ions

Result 1:

THEMIS-ARTEMIS observations of demagnetized ions and magnetized electrons



Agreement between EFI E-field data and plasma derived hall field.

Bx_sse
By_sse
Bz_sse

Ion Energy Flux
[eV/eV cm² sr s]

Electron Energy Flux
[eV/eV cm² sr s]

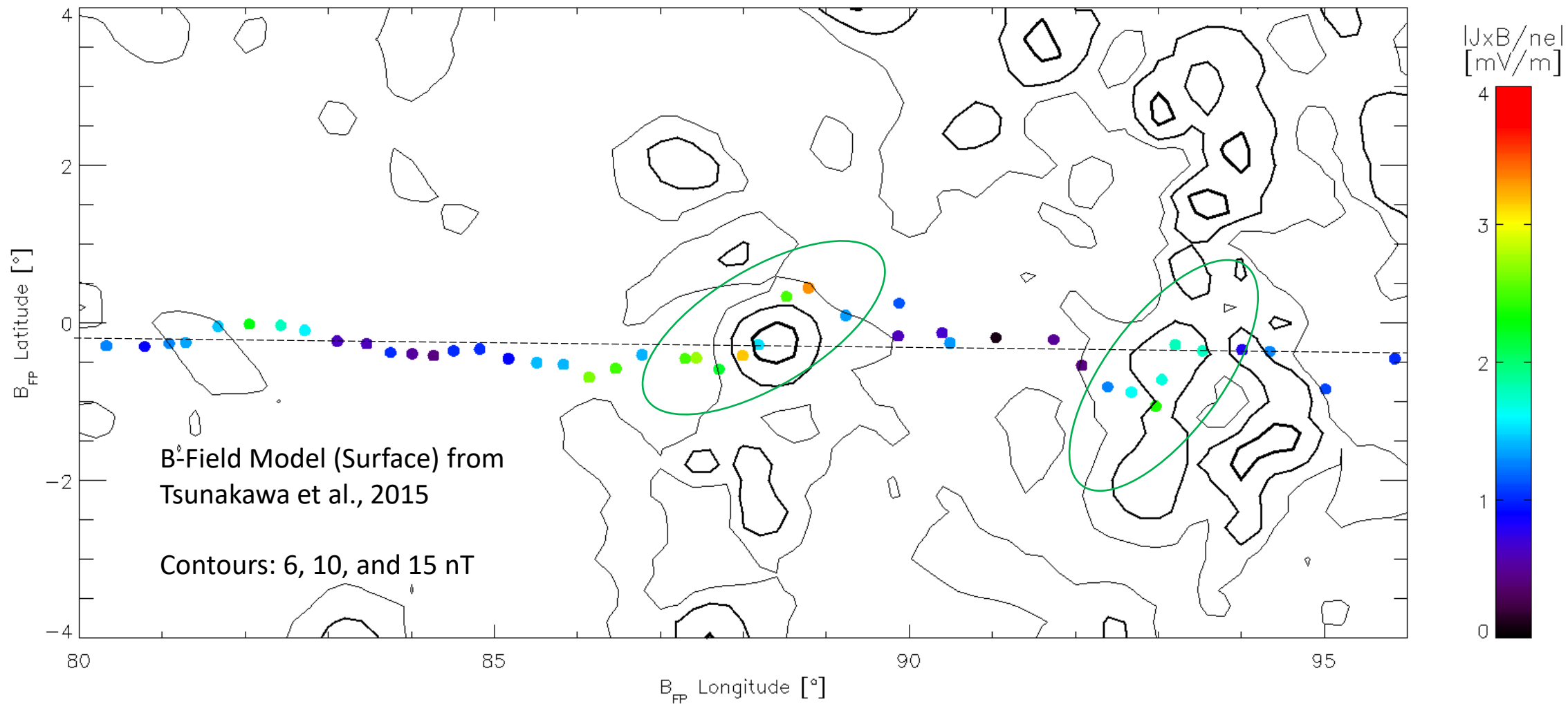
Electron Energy Flux
[eV/eV cm² sr s]

V_{e,z,sse}
V_{i,z,sse}
E_{0xB_{z,sse}}

(JxB/ne_Plasma)_{x,sse}
(E + V_ixB)_{x,sse}

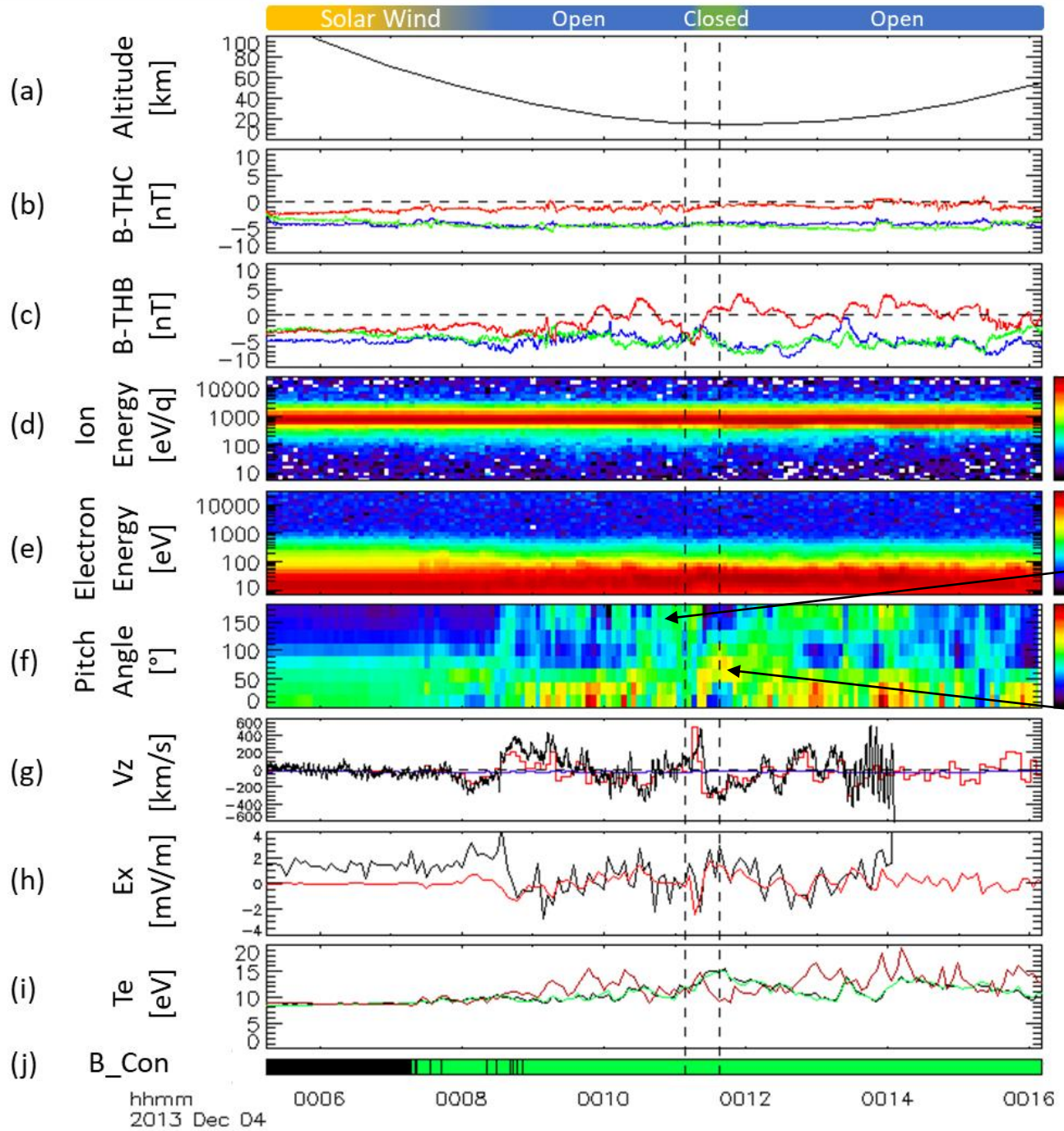
Te_para
Te_perp1
Te_perp2

hhmm
2013 Dec 04
0006 0008 0010 0012 0014 0016



Localized enhancement of Hall E-Field coincident
with crustal magnetization.

Result 2: THEMIS-ARTEMIS observations of field line topology change



Counter-streaming electron beams
indicative of “open” field line
topology

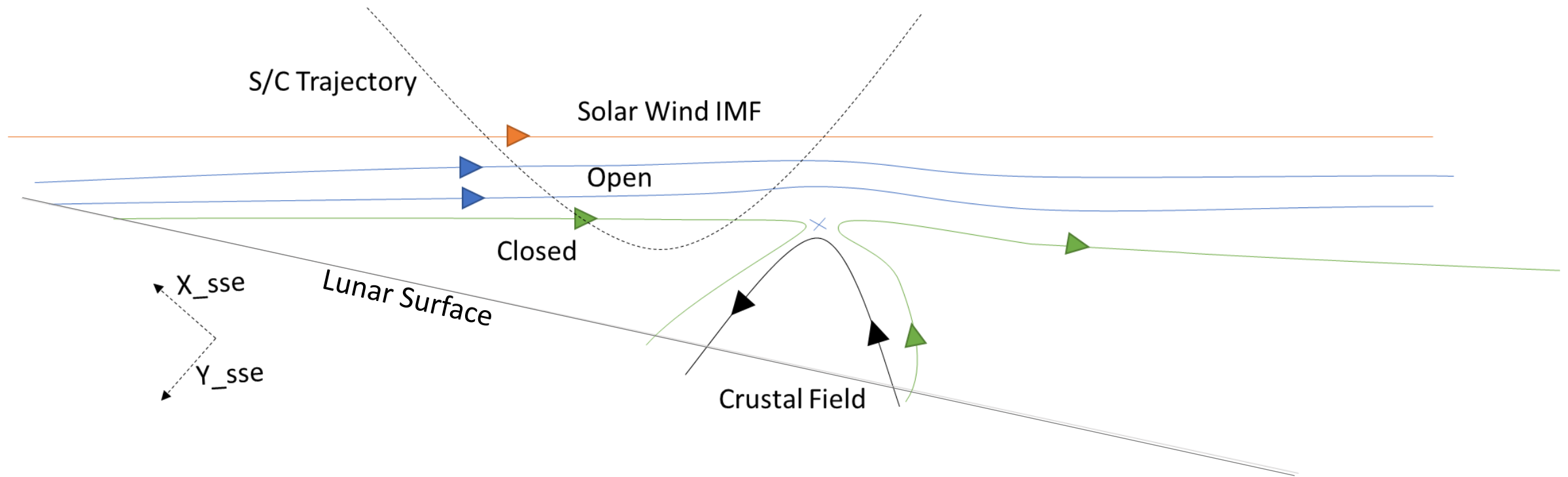
Double-sided loss cone distribution
indicative of a “closed” field line
topology

Change in topology driven by
reconnection between the IMF and
crustal magnetic fields?

- Bx_sse
- By_sse
- Bz_sse
- Bx_sse
- By_sse
- Bz_sse
- Ion Energy Flux [eV/eV cm² sr s]
- Electron Energy Flux [eV/eV cm² sr s]
- Electron Energy Flux [eV/eV cm² sr s]
- V_{e,z,sse}
- V_{i,z,sse}
- E₀xB_{z,sse}
- (JxB/ne_{Plasma})_{x,sse}
- (E + V_ixB)_{x,sse}
- Te_para
- Te_perp1
- Te_perp2

Result 2:

Reconnection may be Contributing to the Closed Field Line Topology Observations



Solar wind IMF reconnected with a crustal field leading to a closed field line topology.

In Conclusion...

- THEMIS-ARTEMIS observations confirm that ions become demagnetized near crustal magnetic fields.
 - The generated Hall E-field is concentrated near regions of crustal magnetization.
- THEMIS-ARTEMIS electron pitch angle distributions suggest that within this region a field line topology change was traversed.
 - The change in field line topology was shown as a transition from counter-streaming electron beams to a double-sided loss cone distribution.
 - The change in field line topology may have been facilitated by the reconnection between IMF and crustal magnetic fields, most likely through electron-only reconnection.

Thank you for your time!

Questions?